

Package ‘CoSMoS’

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Type Package

Title Complete Stochastic Modelling Solution

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Description A single framework, unifying, extending, and improving a general-purpose modelling strategy, based on the assumption that any process can emerge by transforming a specific 'parent' Gaussian process Papalexiou (2018) <doi:10.1016/j.advwatres.2018.02.013>.

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CoSMoS-package	<i>CoSMoS: Complete Stochastic Modelling Solution</i>
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Description

CoSMoS is an R package that makes time series generation with desired properties easy. Just choose the characteristics of the time series you want to generate, and it will do the rest.

Details

The generated time series preserve any probability distribution and any linear autocorrelation structure. Users can generate as many and as long time series from processes such as precipitation, wind, temperature, relative humidity etc. It is based on a framework that unified, extended, and improved a modelling strategy that generates time series by transforming “parent” Gaussian time series having specific characteristics (Papalexiou, 2018).

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References

Papalexiou, S.M., 2018. Unified theory for stochastic modelling of hydroclimatic processes: Preserving marginal distributions, correlation structures, and intermittency. *Advances in Water Resources* 115, 234-252. ([link](#))

Description

Provides a parametric function that describes the values of the linear autocorrelation up to desired lags. For more details on the parametric autocorrelation structures see section 3.2 in [\(Papalexiou 2018\)](#)

Usage

```
acs(id, ...)
```

Arguments

id	autocorrelation structure id
...	other arguments (t as lag and acs parameters)

Examples

```
library(CoSMoS); library(ggplot2)

## specify lag
t <- 0:10

## get the ACS
f <- acs('fgn', t = t, H = .75)
b <- acs('burrXII', t = t, scale = 1, shape1 = .6, shape2 = .4)
w <- acs('weibull', t = t, scale = 2, shape = 0.8)
p <- acs('paretoII', t = t, scale = 3, shape = 0.3)

## visualize the ACS
dta <- data.frame(t, f, b, w, p)

m.dta <- melt(dta, id.vars = 't')

ggplot(m.dta,
       aes(x = t,
           y = value,
           group = variable,
           colour = variable)) +
  geom_point(size = 2.5) +
  geom_line(lwd = 1) +
  scale_color_manual(values = c('steelblue4', 'red4', 'green4', 'darkorange'),
                    labels = c('FGN', 'Burr XII', 'Weibull', 'Pareto II'),
                    name = '') +
  labs(x = bquote(lag ~ tau),
       y = 'Acf') +
```

```
scale_x_continuous(breaks = t) +
theme_classic()
```

actpnts *AutoCorrelation Transformed Points*

Description

Transforms a gaussian process in order to match a target marginal lowers its autocorrelation values. The actpnts evaluates the corresponding autocorrelations for the given target marginal for a set of gaussian correlations, i.e., it returns (ρ_x, ρ_z) points where ρ_x and ρ_z represent, respectively, the autocorrelations of the target and gaussian process.

Usage

```
actpnts(margdist, margarg, p0 = 0, distbounds = c(-Inf, Inf))
```

Arguments

margdist	target marginal distribution
margarg	list of marginal distribution arguments
p0	probability zero
distbounds	distribution bounds (default set to c(-Inf, Inf))

Examples

```
library(CoSMoS); library(ggplot2)

## here we target to a process that has the Pareto type II marginal distribution
## with scale parameter 1 and shape parameter 0.3
## (note that all parameters have to be named)
dist <- 'paretoII'
distarg <- list(scale = 1, shape = .3)

x <- actpnts(margdist = dist, margarg = distarg, p0 = 0)
x

## you can see the points by using
ggplot(x,
  aes(x = rhox,
      y = rhoz)) +
  geom_point(colour = 'royalblue4', size = 2.5) +
  geom_abline(lty = 5) +
  labs(x = bquote(Autocorrelation ~ rho[x]),
      y = bquote(Gaussian ~ rho[z])) +
  scale_x_continuous(limits = c(0, 1)) +
```

```
scale_y_continuous(limits = c(0, 1)) +
theme_classic()
```

analyzeTS

The Functions analyzeTS, reportTS, and simulateTS

Description

Provide a complete set of tools to make time series analysis a piece of cake - analyzeTS() automatically performs seasonal analysis, fits distributions and correlation structures, reportTS provides visualizations of the fitted distributions and correlation structures, and a table with the values of the fitted parameters and basic descriptive statistics, simulateTS automatically takes the results of the analyzeTS and generates syntetic ones.

Usage

```
analyzeTS(TS, season = "month", dist = "ggamma", acsID = "weibull",
  norm = "N2", n.points = 30, lag.max = 30, constrain = FALSE)
```

```
reportTS(aTS, method = "dist")
```

```
simulateTS(aTS, from = NULL, to = NULL)
```

Arguments

TS	time series in format - date, value
season	name of the season (e.g. month, week)
dist	name of the distribution to be fitted
acsID	ID of the autocorrelation structure to be fitted
norm	norm used for distribution fitting - id ('N1', 'N2', 'N3', 'N4')
n.points	number of points to be subsetted from ecdf
lag.max	max lag for the empirical autocorrelation structure
constrain	logical - constrain shape2 parametes for finite tails
aTS	analyzed timeseries
method	report method - 'dist' for distribution fits, 'acs' for ACS fits and 'stat' for basic statistical report
from	starting date/time of the simulation
to	end date/time of the simulation

Details

In practice, we usually want to simulate a natural process using some sampled time series. To generate a synthetic time series with similar characteristics to the observed values, we have to determine marginal distribution, autocorrelation structure and probability zero for each individual month. This can be done by fitting distributions and autocorrelation structures with `analyzeTS()`. Result can be checked with `reportTS()`. Synthetic time series with the same statistical properties can be produced with `simulateTS()`.

Recommended distributions for variables:

- *precipitation*: `ggamma` (Generalized Gamma), `burr###` (Burr type)
- *streamflow*: `ggamma` (Generalized Gamma), `burr###` (Burr type)
- *relative humidity*: `beta`
- *temperature*: `norm` (Normal distribution)

Examples

```
library(CoSMoS)

## Load data included in the package
## (to find out more about the data use ?precip)
data('precip')

## Fit seasonal ACSs and distributions to the data
a <- analyzeTS(precip)

reportTS(a, 'dist') ## show seasonal distribution fit
reportTS(a, 'acs') ## show seasonal ACS fit
reportTS(a, 'stat') ## display basic descriptive statistics

#####
## 'duplicate' analyzed time series ##
sim <- simulateTS(a)

## plot the result
precip[, id := 'observed']
sim[, id := 'simulated']

dta <- rbind(precip, sim)

ggplot(dta) +
  geom_line(aes(x = date, y = value)) +
  facet_wrap(~id, ncol = 1) +
  theme_classic()

#####
## or simulate timeseries of different length ##
sim <- simulateTS(a,
  from = as.POSIXct('1978-12-01 00:00:00'),
  to = as.POSIXct('2008-12-01 00:00:00'))
```

```
## and plot the result
precip[, id := 'observed']
sim[, id := 'simulated']

dta <- rbind(precip, sim)

ggplot(dta) +
  geom_line(aes(x = date, y = value)) +
  facet_wrap(~id, ncol = 1) +
  theme_classic()
```

BurrIII

*Burr Type III distribution***Description**

Provides density, distribution function, quantile function, random value generation, and raw moments of order r for the Burr Type III distribution.

Usage

```
dburrIII(x, scale, shape1, shape2, log = FALSE)

pburrIII(q, scale, shape1, shape2, lower.tail = TRUE, log.p = FALSE)

qburrIII(p, scale, shape1, shape2, lower.tail = TRUE, log.p = FALSE)

rburrIII(n, scale, shape1, shape2)

mburrIII(r, scale, shape1, shape2)
```

Arguments

<code>x, q</code>	vector of quantiles.
<code>scale, shape1, shape2</code>	scale and shape parameters; the shape arguments cannot be a vectors (must have length one).
<code>log, log.p</code>	logical; if TRUE, probabilities p are given as $\log(p)$.
<code>lower.tail</code>	logical; if TRUE (default), probabilities are $P[X \leq x]$ otherwise, $P[X > x]$.
<code>p</code>	vector of probabilities.
<code>n</code>	number of observations. If $\text{length}(n) > 1$, the length is taken to be the number required.
<code>r</code>	raw moment order

Examples

```
## plot the density

ggplot(data.frame(x = c(1, 15)),
  aes(x)) +
  stat_function(fun = dburrIII,
    args = list(scale = 5,
      shape1 = .25,
      shape2 = .75),
    colour = 'royalblue4') +
  labs(x = '',
    y = 'Density') +
  theme_classic()
```

BurrXII

*Burr Type XII distribution***Description**

Provides density, distribution function, quantile function, random value generation, and raw moments of order r for the Burr Type XII distribution.

Usage

```
dburrXII(x, scale, shape1, shape2, log = FALSE)

pburrXII(q, scale, shape1, shape2, lower.tail = TRUE, log.p = FALSE)

qburrXII(p, scale, shape1, shape2, lower.tail = TRUE, log.p = FALSE)

rburrXII(n, scale, shape1, shape2)

mburrXII(r, scale, shape1, shape2)
```

Arguments

<code>x, q</code>	vector of quantiles.
<code>scale, shape1, shape2</code>	scale and shape parameters; the shape arguments cannot be a vectors (must have length one).
<code>log, log.p</code>	logical; if TRUE, probabilities p are given as $\log(p)$.
<code>lower.tail</code>	logical; if TRUE (default), probabilities are $P[X \leq x]$ otherwise, $P[X > x]$.
<code>p</code>	vector of probabilities.
<code>n</code>	number of observations. If $\text{length}(n) > 1$, the length is taken to be the number required.
<code>r</code>	raw moment order

Examples

```
## plot the density

ggplot(data.frame(x = c(0, 10)),
  aes(x)) +
  stat_function(fun = dburrXII,
    args = list(scale = 5,
      shape1 = .25,
      shape2 = .75),
    colour = 'royalblue4') +
  labs(x = '',
    y = 'Density') +
  theme_classic()
```

 checkTS

Check generated timeseries

Description

Compares generated time series sample statistics with the theoretically expected values.

Usage

```
checkTS(TS, distbounds = c(-Inf, Inf))
```

Arguments

TS	generated timeseries
distbounds	distribution bounds (default set to c(-Inf, Inf))

Examples

```
library(CoSMoS)

## check your generated timeseries
x <- generateTS(margdist = 'burrXII',
  margarg = list(scale = 1,
    shape1 = .75,
    shape2 = .25),
  acsvalue = acs(id = 'weibull',
    t = 0:30,
    scale = 10,
    shape = .75),
  n = 1000, p = 30, p0 = .5, TSn = 5)

checkTS(x)
```

 fitactf

Fit the AutoCorrelation Transformation Function

Description

Fits the ACTF (Autocorrelation Transformation Function) to the estimated points (ρ_x, ρ_z) using nls

Usage

```
fitactf(actpnts, discrete = FALSE)
```

Arguments

actpnts	estimated ACT points
discrete	logical - is the marginal distribution discrete?

Examples

```
library(CoSMoS)

## choose the marginal distribution as Pareto type II with corresponding parameters
dist <- 'paretoII'
distarg <- list(scale = 1, shape = .3)

## estimate rho 'x' and 'z' points using ACTI
p <- actpnts(margdist = dist, margarg = distarg, p0 = 0)

## fit ACTF
fit <- fitactf(p)

## plot the result
plot(fit, method = 'base')
```

 fitDist

Distribution fitting

Description

Uses Nelder-Mead simplex algorithm to minimize fitting norms.

Usage

```
fitDist(data, dist, n.points, norm, constrain)
```

Arguments

data	value to be fitted
dist	name of the distribution to be fitted
n.points	number of points to be subsetted from ecdf
norm	norm used for distribution fitting - id ('N1', 'N2', 'N3', 'N4')
constrain	logical - constrain shape2 parametes for finite tails

Examples

```
x <- fitDist(rnorm(1000), 'norm', 30, 'N1', FALSE)
x
```

generateTS

Generate timeseries

Description

Generates timeseries with given properties, just provide (1) the target marginal distribution and its parameters, (2) the target autocorrelation structure or individual autocorrelation values up to a desired lag, and (3) the probability zero if you wish to simulate an intermittent process.

Usage

```
generateTS(n, margdist, margarg, p = NULL, p0 = 0, TSn = 1,
  distbounds = c(-Inf, Inf), acsvalue = NULL)
```

Arguments

n	number of values
margdist	target marginal distribution
margarg	list of marginal distribution arguments
p	integer - model order (if NULL - limits maximum model order according to auto-correlation structure values)
p0	probability zero
TSn	number of timeseries to be generated
distbounds	distribution bounds (default set to c(-Inf, Inf))
acsvalue	target auto-correlation structure (from lag 0)

Details

A step-by-step guide:

- First define the target marginal (margdist), that is, the probability distribution of the generated data. For example set margdist = 'ggamma' if you wish to generate data following the Generalized Gamma distribution, margdist = 'burrXII' for Burr type XII distribution etc. For a full list of the distributions we support see the help vignette: `vignette('vignette', package = 'CoSMoS')`. In general, the package supports all build-in distribution functions of R and of other packages.
- Define the parameters' values (margarg) of the distribution you selected. For example the Generalized Gamma has one scale and two shape parameters so set the desired value, e.g., `margarg = list(scale = 2, shape1 = 0.9, shape2 = 0.8)`. Note distributions might have different number of parameters and different type of parameters (location, scale, shape). To see the parameters of each distribution we support, see the help vignette: `vignette('vignette', package = 'CoSMoS')`.
- If you wish your time series to be intermittent (e.g., precipitation), then define the probability zero. For example, set `p0 = 0.9`, if you wish your generated data to have 90% of zero values (dry days).
- Define your linear autocorrelations.
 - You can supply specific lag autocorrelations starting from lag 0 and up to a desired lag, e.g., `acs = c(1, 0.9, 0.8, 0.7)`; this will generate a process with lag1, 2 and 3 autocorrelations equal with 0.9, 0.8 and 0.7.
 - Alternatively, you can use a parametric autocorrelation structure (see section 3.2 in [Papalexiou 2018](#)). We support the following autocorrelation structures (acs) weibull, paretoII, fgn and burrXII. See also [acs](#) examples.
- Define the order to the autoregressive model p. For example if you aim to preserve the first 10 lag autocorrelations then just set `p = 10`. Otherwise set it `p = NULL` and the model will decide the value of p in order to preserve the whole autocorrelation structure.
- Lastly just define the time series length, e.g., `n = 1000` and number of time series you wish to generate, e.g., `TSn = 10`.

Play around with the following given examples which will make the whole process a piece of cake.

Examples

```
library(CoSMoS)

## Case1:
## You wish to generate 3 time series of size 1000 each
## that follow the Generalized Gamma distribution with parameters
## scale = 1, shape1 = 0.8, shape2 = 0.8
## and autocorrelation structure the ParetoII
## with parameters scale = 1 and shape = .75
x <- generateTS(margdist = 'ggamma',
                margarg = list(scale = 1,
                              shape1 = .8,
                              shape2 = .8),
                acsvalue = acs(id = 'paretoII',
                              t = 0:30,
```

```

                                scale = 1,
                                shape = .75),
    n = 1000,
    p = 30,
    TSn = 3)

## see the results
plot(x)

## Case2:
## You wish to generate time series the same distribution
## and autocorrelations as is Case1 but intermittent
## with probability zero equal to 90%
y <- generateTS(margdist = 'ggamma',
                margarg = list(scale = 1,
                              shape1 = .8,
                              shape2 = .8),
                acsvalue = acs(id = 'paretoII',
                              t = 0:30,
                              scale = 1,
                              shape = .75),
                p0 = .9,
                n = 1000,
                p = 30,
                TSn = 3)

## see the results
plot(y)

## Case3:
## You wish to generate a time series of size 1000
## that follows the Beta distribution
## (e.g., relative humidity ranging from 0 to 1)
## with parameters shape1 = 0.8, shape2 = 0.8, is defined from 0 to 1
## and autocorrelation structure the ParetoII
## with parameters scale = 1 and shape = .75
z <- generateTS(margdist = 'beta',
                margarg = list(shape1 = .6,
                              shape2 = .8),
                distbounds = c(0, 1),
                acsvalue = acs(id = 'paretoII',
                              t = 0:30,
                              scale = 1,
                              shape = .75),
                n = 1000,
                p = 20)

## see the results
plot(z)

## Case4:
```

```
## Same in previous case but now you provide specific
## autocorrelation values for the first three lags,
## ie., lag 1 to 3 equal to 0.9, 0.8 and 0.7

z <- generateTS(margdist = 'beta',
               margarg = list(shape1 = .6,
                             shape2 = .8),
               distbounds = c(0, 1),
               acsvalue = c(1, .9, .8, .7),
               n = 1000,
               p = TRUE)

## see the results
plot(z)
```

GGamma

Generalized gamma distribution

Description

Provides density, distribution function, quantile function, random value generation, and raw moments of order r for the generalized gamma distribution.

Usage

```
dggamma(x, scale, shape1, shape2, log = FALSE)
```

```
pggamma(q, scale, shape1, shape2, lower.tail = TRUE, log.p = FALSE)
```

```
qggamma(p, scale, shape1, shape2, lower.tail = TRUE, log.p = FALSE)
```

```
rggamma(n, scale, shape1, shape2)
```

```
mgamma(r, scale, shape1, shape2)
```

Arguments

<code>x, q</code>	vector of quantiles.
<code>scale, shape1, shape2</code>	scale and shape parameters; the shape arguments cannot be a vectors (must have length one).
<code>log, log.p</code>	logical; if TRUE, probabilities p are given as $\log(p)$.
<code>lower.tail</code>	logical; if TRUE (default), probabilities are $P[X \leq x]$ otherwise, $P[X > x]$.
<code>p</code>	vector of probabilities.

n	number of observations. If <code>length(n) > 1</code> , the length is taken to be the number required.
r	raw moment order

Examples

```
## plot the density

ggplot(data.frame(x = c(0, 20)),
       aes(x)) +
  stat_function(fun = dgamma,
              args = list(scale = 5,
                          shape1 = .25,
                          shape2 = .75),
              colour = 'royalblue4') +
  labs(x = '',
       y = 'Density') +
  theme_classic()
```

moments

Numerical estimation of moments

Description

Uses numerical integration to calculate the theoretical raw or central moments of the specified distribution

Usage

```
moments(dist, distarg, p0 = 0, raw = T, central = T, coef = T,
        distbounds = c(-Inf, Inf), order = 1:4)
```

Arguments

dist	distribution
distarg	list of distribution arguments
p0	probability zero
raw	logical - calculate raw moments?
central	logical - calculate central moments?
coef	logical - calculate coefficients (coefficient of variation, skewness and kurtosis)?
distbounds	distribution bounds (default set to <code>c(-Inf, Inf)</code>)
order	vector of integers - raw moment orders

Examples

```

library(CoSMoS)

## Normal Distribution
moments('norm', list(mean = 2, sd = 1))

## Pareto type II
scale <- 1
shape <- .2

moments(dist = 'paretoII',
        distarg = list(shape = shape,
                       scale = scale))

```

ParetoII

Pareto type II distribution

Description

Provides density, distribution function, quantile function, random value generation and raw moments of order r for the Pareto type II distribution.

Usage

```

dparetoII(x, scale, shape, log = FALSE)

pparetoII(q, scale, shape, lower.tail = TRUE, log.p = FALSE)

qparetoII(p, scale, shape, lower.tail = TRUE, log.p = FALSE)

rparetoII(n, scale, shape)

mparetoII(r, scale, shape)

```

Arguments

<code>x, q</code>	vector of quantiles.
<code>scale, shape</code>	scale and shape parameters; the shape argument cannot be a vector (must have length one).
<code>log, log.p</code>	logical; if TRUE, probabilities p are given as $\log(p)$.
<code>lower.tail</code>	logical; if TRUE (default), probabilities are $P[X \leq x]$ otherwise, $P[X > x]$.
<code>p</code>	vector of probabilities.
<code>n</code>	number of observations. If $\text{length}(n) > 1$, the length is taken to be the number required.
<code>r</code>	raw moment order

Examples

```
## plot the density

ggplot(data.frame(x = c(0, 20)),
  aes(x)) +
  stat_function(fun = dparetoII,
    args = list(scale = 1,
      shape = .3),
    colour = 'royalblue4') +
  labs(x = '',
    y = 'Density') +
  theme_classic()
```

plot.acti

AutoCorrelation Transformation Function visualisation

Description

Visualizes the autocorrelation transformation integral (there are two possible methods for plotting - base graphics and ggplot2 package)

Usage

```
## S3 method for class 'acti'
plot(x, ...)
```

Arguments

x	fitactf result object
...	other arguments

Examples

```
library(CoSMoS)

## choose the marginal distribution as Pareto type II with corresponding parameters
dist <- 'paretoII'
distarg <- list(scale = 1, shape = .3)

## estimate rho 'x' and 'z' points using ACTI
p <- actpnts(margdist = dist, margarg = distarg, p0 = 0)

## fit ACTF
fit <- fitactf(p)

## plot the results
```

```
plot(fit)
plot(fit, main = 'Pareto type II distribution \nautocorrelation tranformation')
```

plot.checkTS *Plot method for check results*

Description

Plot method for check results

Usage

```
## S3 method for class 'checkTS'
plot(x, ...)
```

Arguments

x	check result
...	other args

Examples

```
library(CoSMoS)

## check your generated timeseries
x <- generateTS(margdist = 'burrXII',
               margarg = list(scale = 1,
                              shape1 = .75,
                              shape2 = .15),
               acsvalue = acs(id = 'weibull',
                              t = 0:30,
                              scale = 10,
                              shape = .75),
               n = 10000, p = 30, p0 = .25, TSn = 100)

chck <- checkTS(x)

plot(chck)
```

plot.cosmosts *Plot generated Timeseries*

Description

Visualizes Timeseries generated by the package CoSMoS

Usage

```
## S3 method for class 'cosmosts'  
plot(x, ...)
```

Arguments

x	fitactf result object
...	other arguments

Examples

```
library(CoSMoS)  
  
## generate TS  
ts <- generateTS(margdist = 'ggamma',  
                  margarg = list(scale = 1,  
                                  shape1 = .8,  
                                  shape2 = .8),  
                  acsvalue = acs(id = 'paretoII',  
                                  t = 0:30,  
                                  scale = 1,  
                                  shape = .75),  
                  n = 1000,  
                  p = 30,  
                  TSn = 2)  
  
## plot the TS  
plot(ts)
```

precip *Hourly station precipitation data*

Description

Station details

- Name: Philadelphia International Airport
- Network ID: COOP:366889
- Latitude/Longitude: 39.87327°, -75.22678°
- Elevation: 3m

Usage

```
precip
```

Format

A data.table with 79633 rows and 2 variables:

date POSIXct format date/time

value precipitation totals

Details

more details can be found [here](#).

Source

The National Oceanic and Atmospheric Administration (NOAA)

regenerateTS

Bulk Timeseries generation

Description

Resamples given Timeseries

Usage

```
regenerateTS(ts, TSn = 1)
```

Arguments

ts generated timeseries using ARp

TSn number of timeseries to be (re)generated

Details

You have used the generateTS function and you wish to generate more time series. Instead of re-running the generateTS you can use the regenerateTS Timeseries that used all the parameters calculated by the generateTS function and thus it is faster.

Examples

```

library(CoSMoS)

## define marginal distribution and arguments with target autocorrelation structure
x <- generateTS(margdist = 'burrXII',
               margarg = list(scale = 1,
                              shape1 = .75,
                              shape2 = .25),
               acsvalue = acs(id = 'weibull',
                              t = 0:30,
                              scale = 10,
                              shape = .75),
               n = 1000, p = 30, p0 = .5, TSn = 3)

## generate new values with same parameters
r <- regenerateTS(x)

plot(r)

```

sample.moments

Estimation of sample moments

Description

Estimation of sample moments

Usage

```

sample.moments(x, na.rm = FALSE, raw = T, central = T, coef = T,
              order = 1:4)

```

Arguments

x	a numeric vector of values
na.rm	a logical value indicating whether NA values should be stripped before the computation proceeds
raw	logical - calculate raw moments?
central	logical - calculate central moments?
coef	logical - calculate coefficients (coefficient of variation, skewness and kurtosis)?
order	vector of integers - raw moment orders

Examples

```
library(CoSMS)

x <- rnorm(1000)
sample.moments(x)

y <- rparetoII(1000, 10, .1)
sample.moments(y)
```

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